Acceleration of Light at Earth's Surface Richard Mould Department of Physics and Astronomy, State University of New York,

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General relativity requires that light traveling upward or downward at the earth's surface has an acceleration equal to +2g.

The radial acceleration of an object at the earth's surface in a gravitational field $g = GM/r^2$ is equal to

$$a_{\rm r} = -g(1 - 3\beta_{\rm r}^{2}) \tag{1}$$

where β_r is the radial velocity v_r divided by c. When $\beta_r = 0$, the object will accelerate downward with acceleration g as expected. When the object in question is light, then $\beta_r = \pm 1$ and a_r is equal to $\pm 2g$. If the object is traveling with a radial velocity $\beta_r = \pm 1/\sqrt{3}$, then it will not experience any radial acceleration at all.

We know that the velocity of light above the surface of the earth is greater than c relative to a surface observer, and it is less than c (in a matter free depression) below the surface. So when traveling from below to above the surface, light will accelerate upward; and when traveling from above to below the surface, it will also accelerate upward. Equation 1 gives the quantitative amount of that acceleration.

This equation is only valid when the gravitational potential ϕ is considerably less that c^2 . More precisely, Eq. 1 is given by

$$a_{\rm r} = -\sigma g (1 - 3\beta_{\rm r}^2/\sigma^2) \tag{2}$$

where $\sigma = 1 + 2\phi/c^2$, with $\phi = -GM/r$ [Ref. 1, Eq. 12.37].

Equation 2 is derived from an even more general relationship in contravariant spatial coordinates $x^{i} = (x^{1}, x^{2}, x^{3})$ given by

$$a^{i} = -\Gamma^{i}_{\mu\nu} v^{\mu} v^{\nu} + (v^{i}/c) \Gamma^{4}_{\mu\nu} v^{\mu} v^{\nu}$$
(3)

where $a^{i} = d^{2}x^{i}/dt^{2}$, $v^{i} = dx^{i}/dt$, and $v^{\mu} = (v^{i}, c)$ is not a four-vector. We call $x^{4} = ct = \tau$. Greek indices sum over four coordinates. The acceleration a_{r} in Eq. 1 is a^{1} in Eq. 3. Equation 3 is derived in Appendix D of Ref. 1.

To evaluate this equation for our case the Schwarzschild metric

$$ds^{2} = \sigma^{-1}dr^{2} + r^{2}(d\theta^{2} + \sin^{2}\theta d\phi^{2}) - \sigma d\tau^{2}$$

is used to find the Christoffel symbols

$$\Gamma^{1}_{11} = -GM/\sigma r^{2}c^{2} \qquad \Gamma^{1}_{44} = \sigma GM/r^{2}c^{2}$$

$$\Gamma^{4}_{14} = \Gamma^{4}_{41} = GM/\sigma r^{2}c^{2}$$

Ref. 1. R. A. Mould, "Basic Relativity", Springer 2002